

Modelling Thin Film Growth

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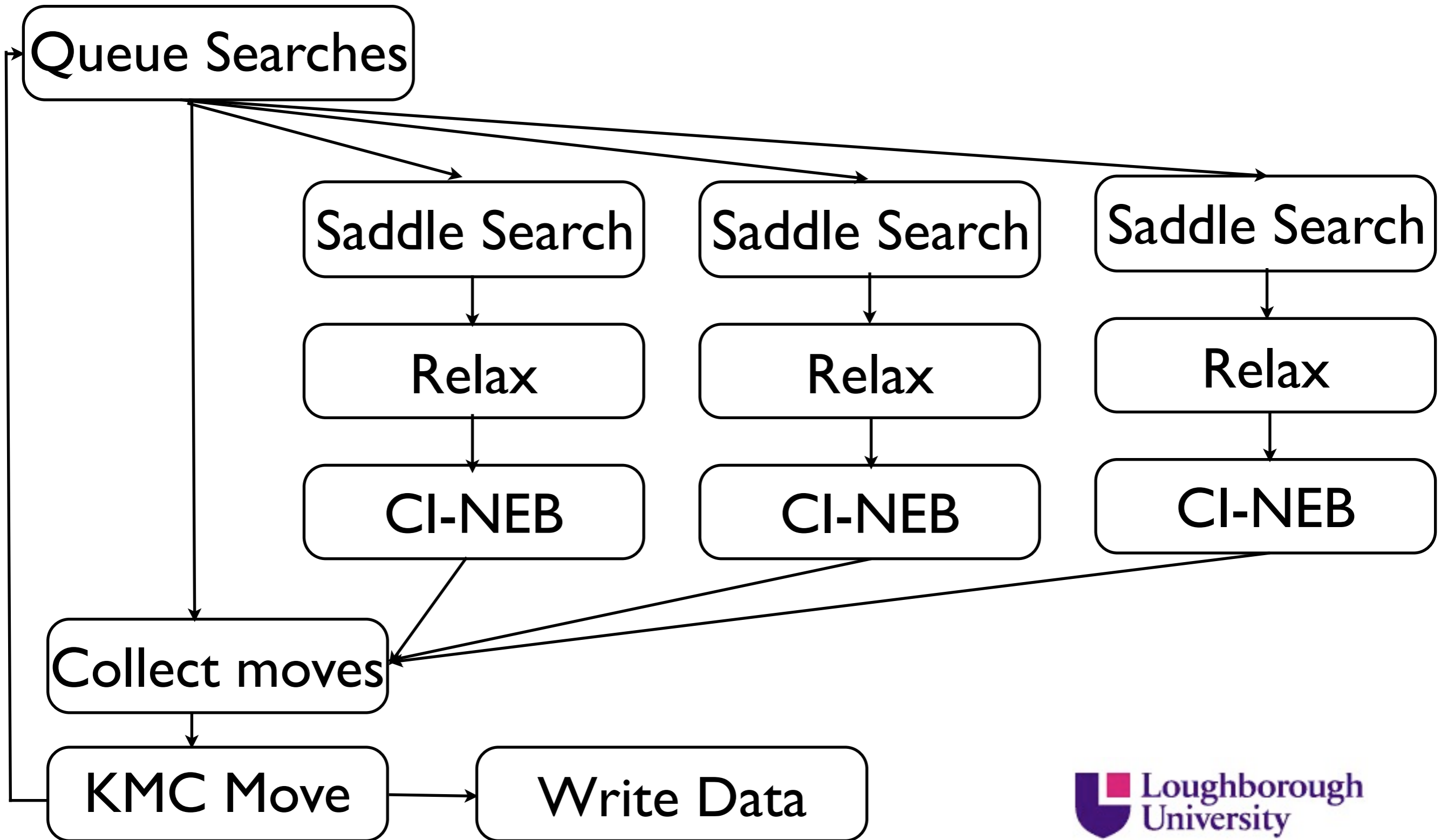
Outline

- Methodology
- Results
- Challenges
- Conclusions

Methodology

- HTST, fixed prefactor
- Mixed MD - otf-kMC
- On-the-fly kinetic Monte Carlo Methodology
 - Defect identification
 - Saddle point search
 - Barrier Calculation
 - KMC Step
 - Repeat

Methodology



Methodology

- Searches are queued and executed when cores are free.
- Mixture of dimer method and RAT used for saddle searches.
- Relaxation performed to find final state.
- CI-NEB used to find barrier height.
- Duplicate saddles discarded.
- Roulette table constructed and move chosen.

Results

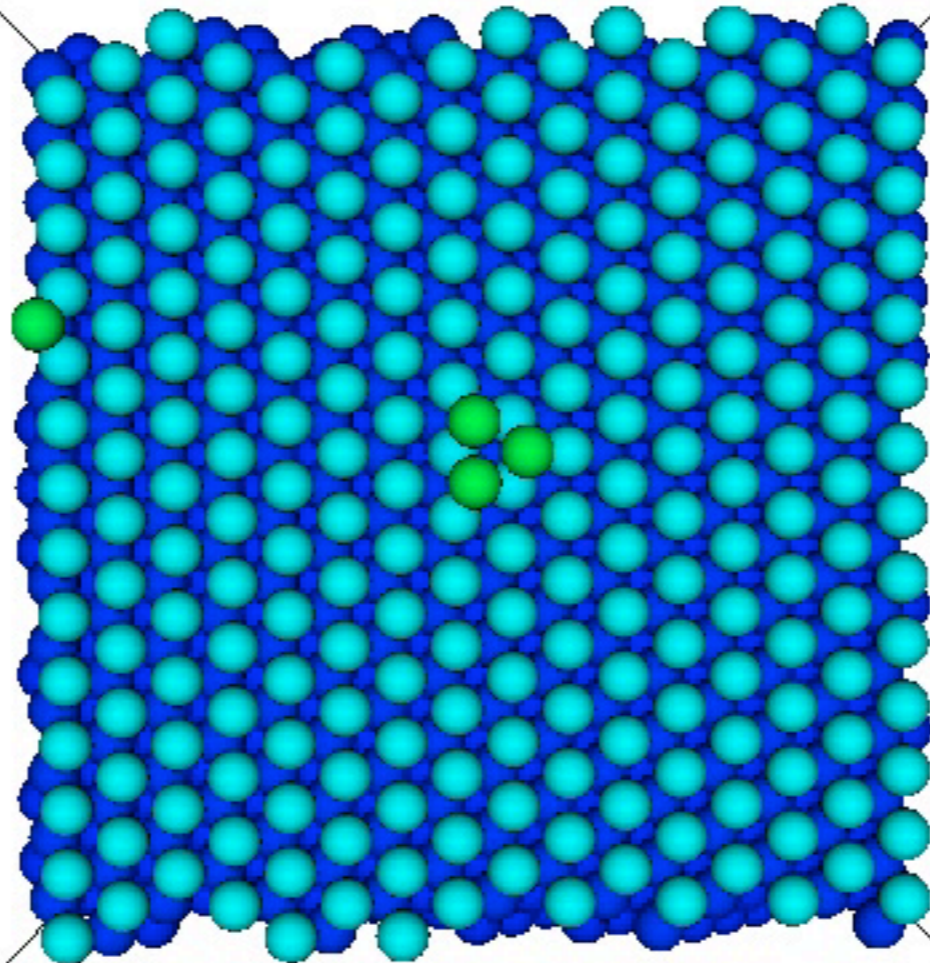
- Simulations of growth of thin films of interest for PV applications.
- Growth of Al and Ag (100) and (111) films - contacts and concentrators.
- Growth of ZnO - TCO.
- Growth of TiO₂ - AR coating.

Growth of Al and Ag

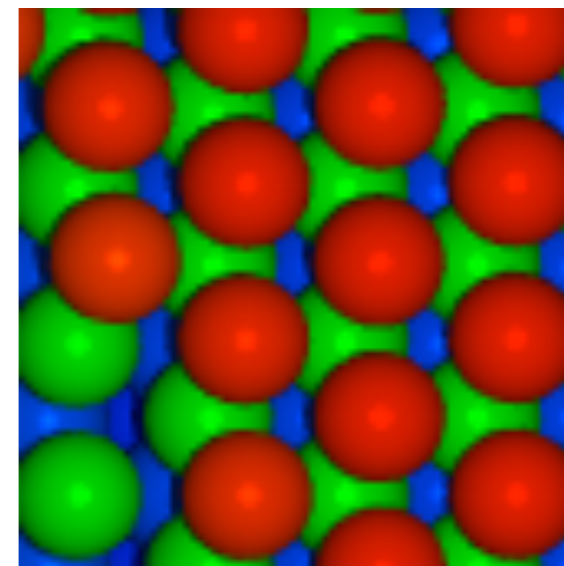
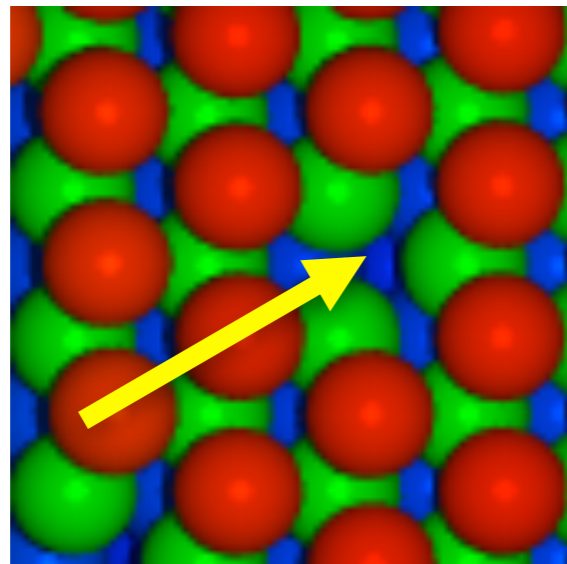
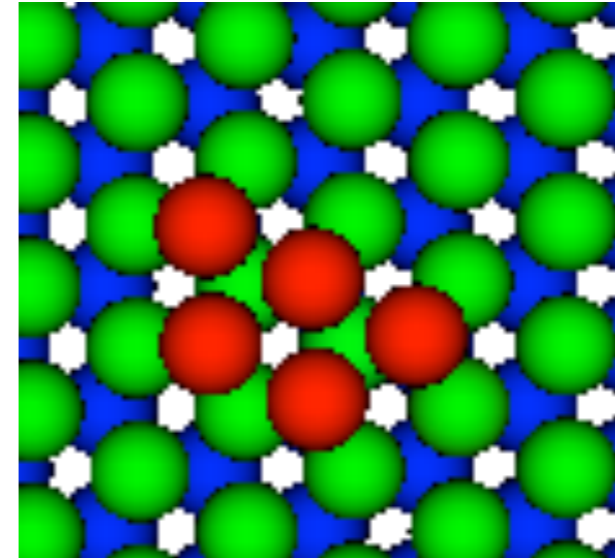
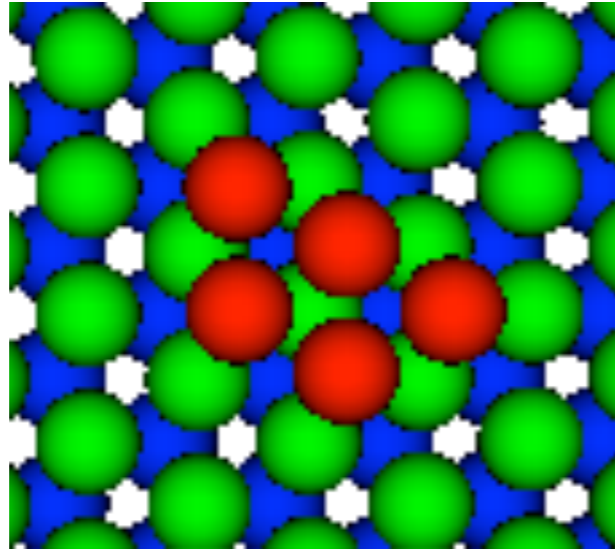
- EAM type potentials.
- 6 initial layers of metal.
- 4 new layers grown.
- 10 ML/s growth rate - 350K temperature.
- Growth energies of 1 eV and 40 eV.
- Optional co-deposition of Ar.

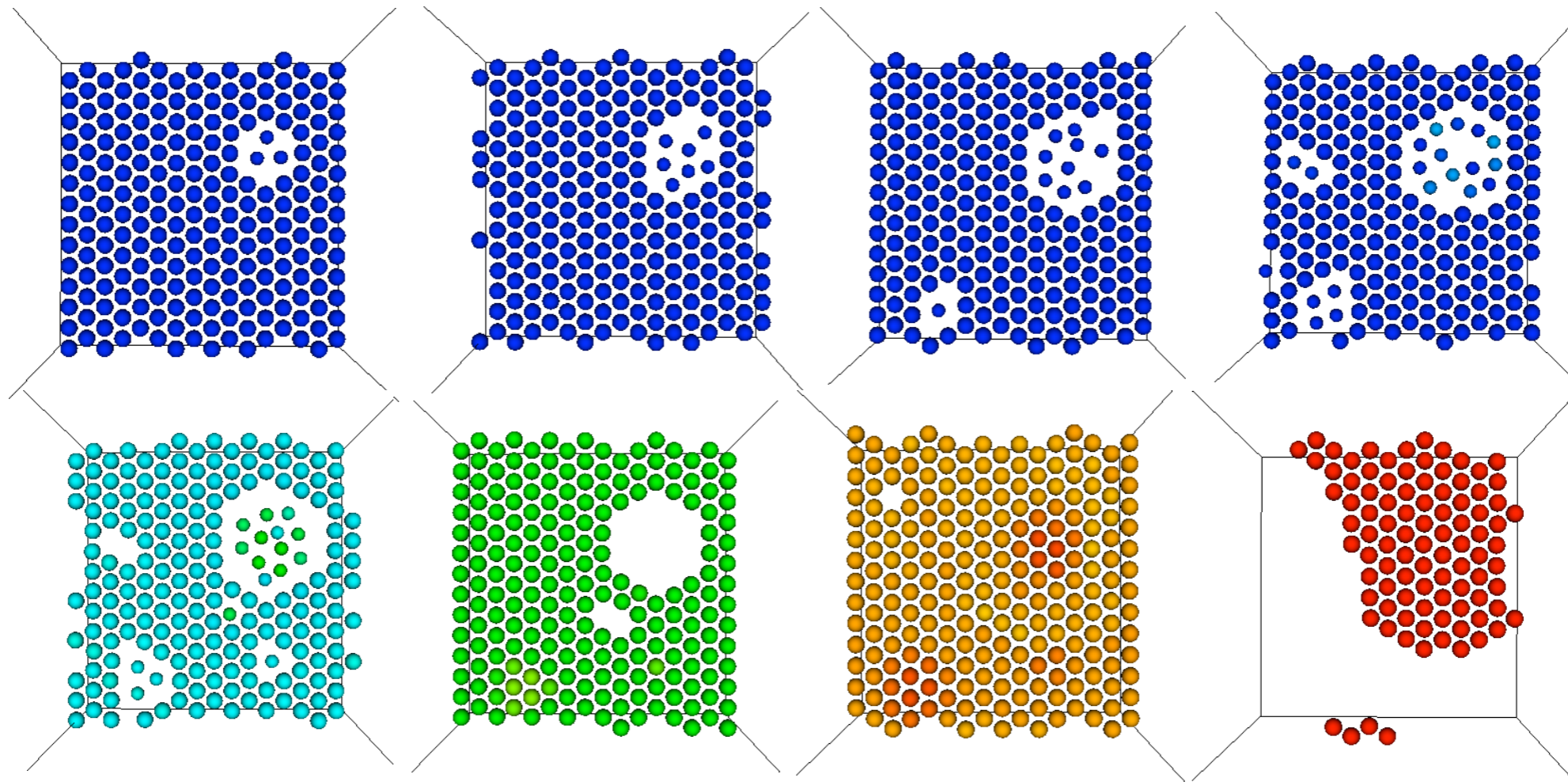
0.000 fs
1348 Atoms
1348 Visible
1348 Al

Al (111) Growth



Concerted Motions



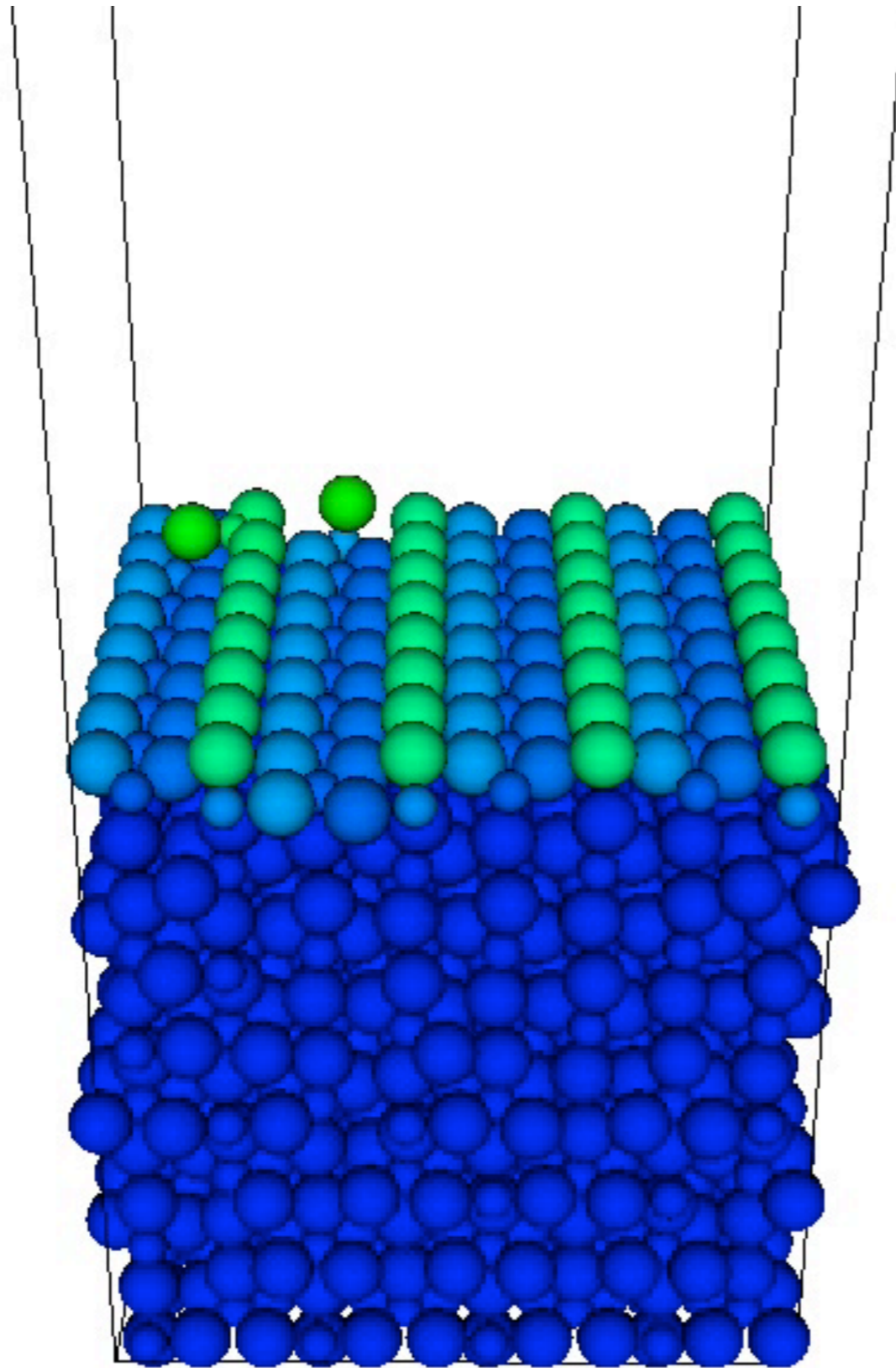


Growth of TiO₂

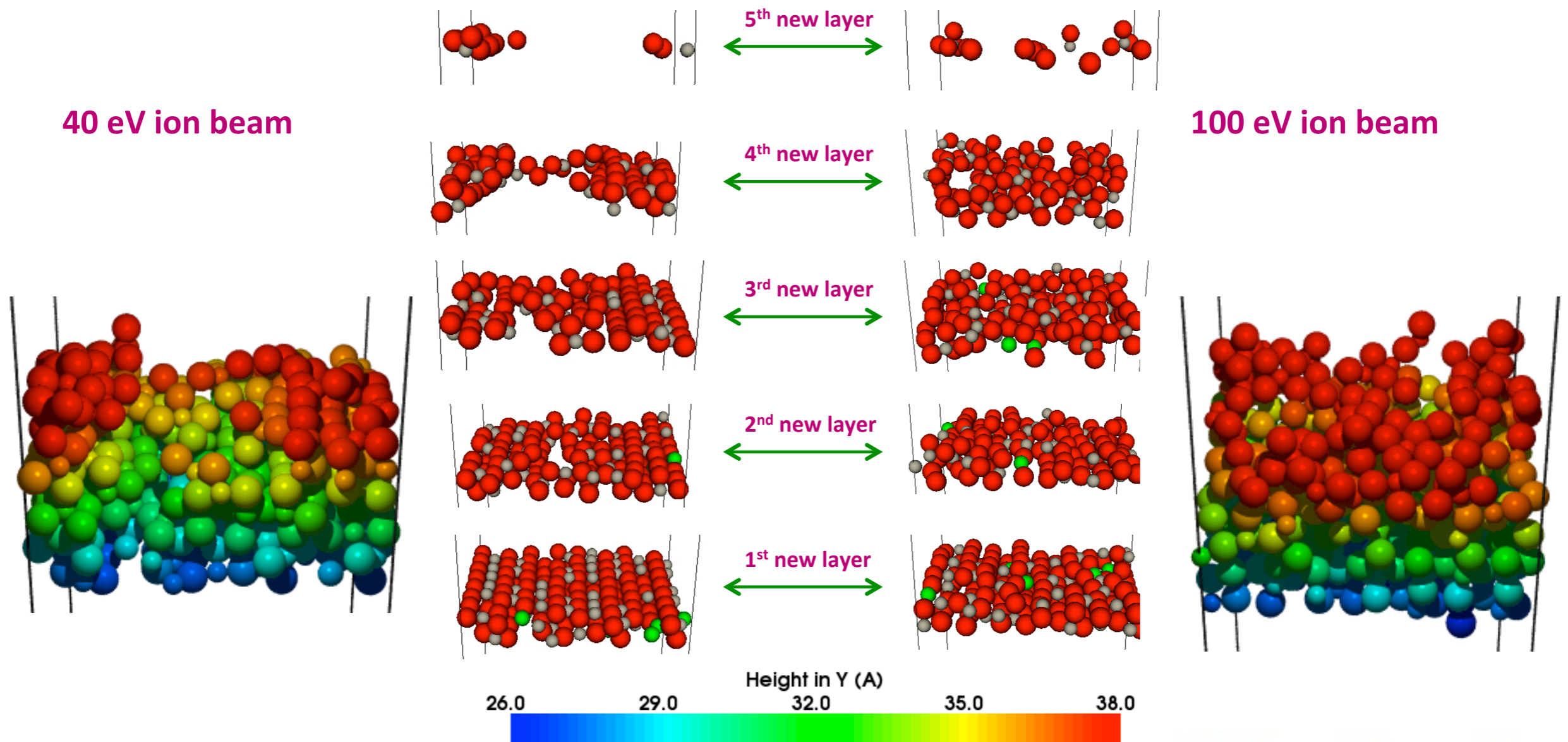
- Variable charge potential used.
- Modified to reproduce important transitions.
- 6 initial layers.
- 4 new layers grown.
- Deposition rate 0.5 ML/s - 350 K.

0.000 fs
1 539 Atoms
1 539 Visible
1 026 O
513 Ti

TiO₂ Growth



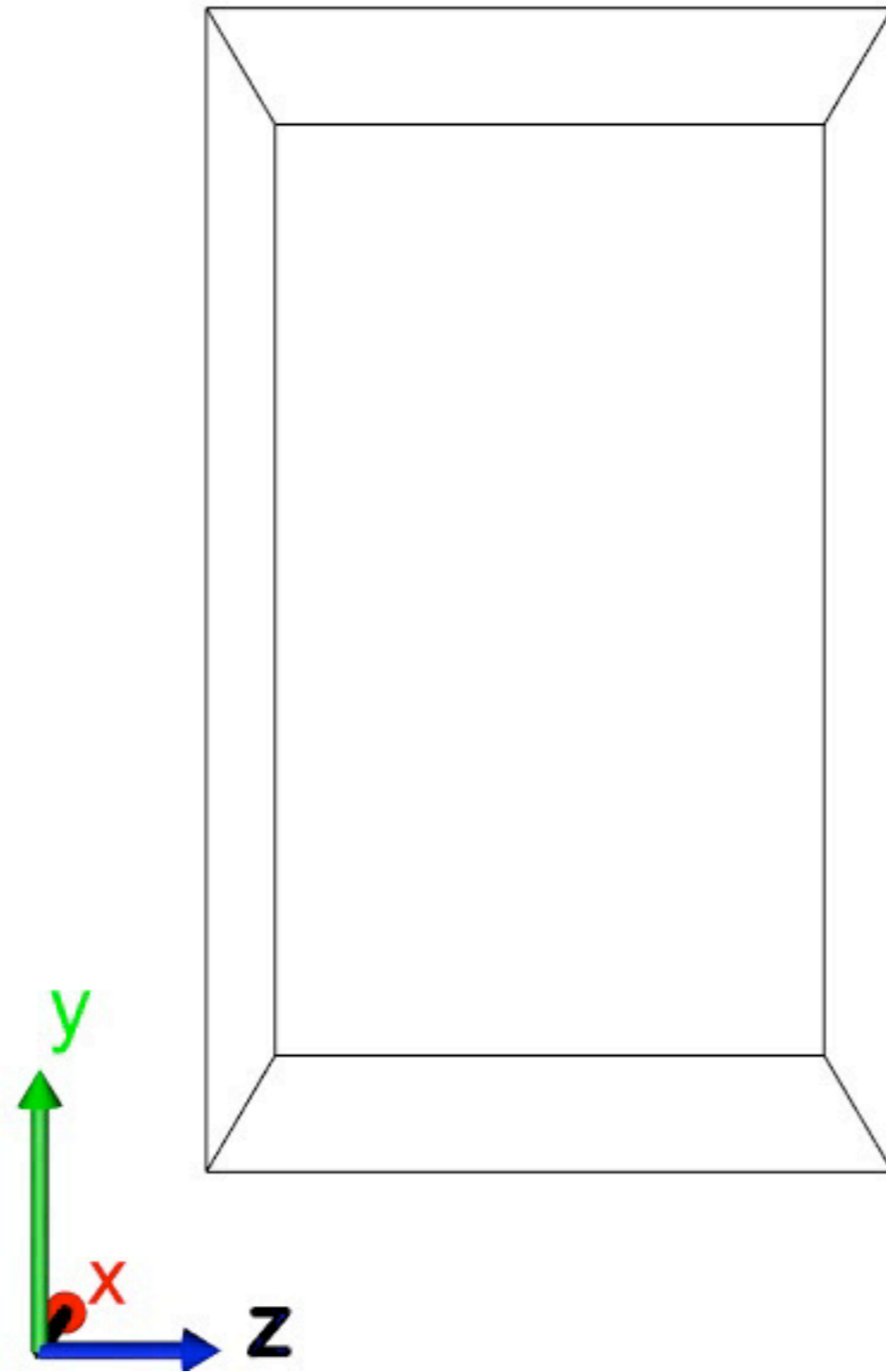
TiO₂ Growth



Multiple Collision Cascades

- 1 keV collision cascade modelled by MD for 20 ps
- Subsequent diffusion modelled by otf-KMC
- Collision cascades performed every 0.2 s, simulates ion implantation rates
- Total of 3 collision cascades modelled

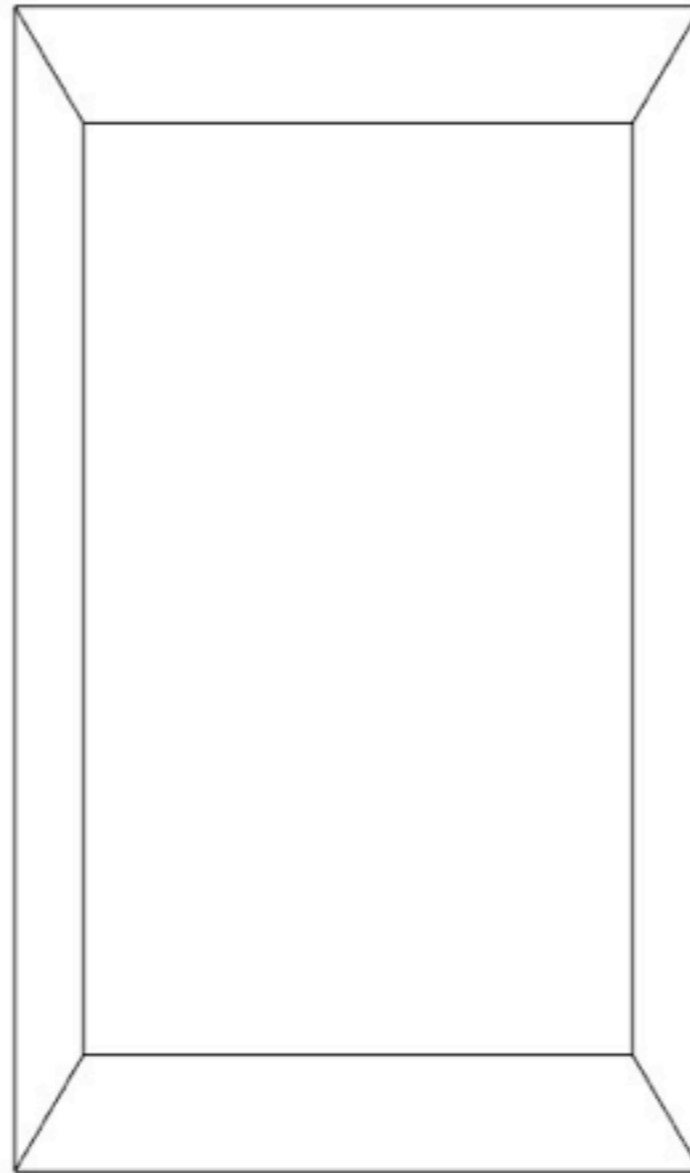
0.00 fs
8470 Atoms
0 Visible



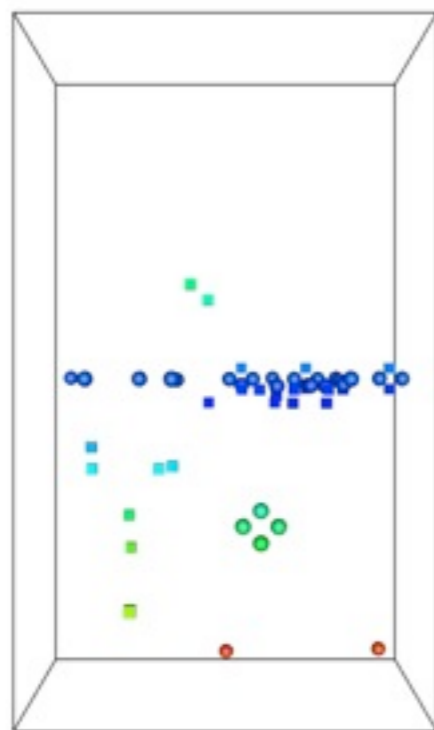
0.00 fs

8470 Atoms

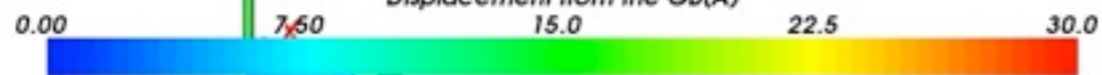
0 Visible



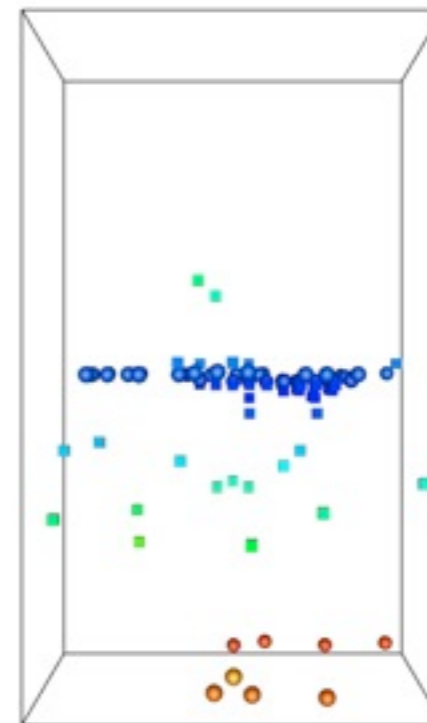
218.96 ms
8470 Atoms
52 Visible
25 Ni Vacancies
27 Ni Interstitials



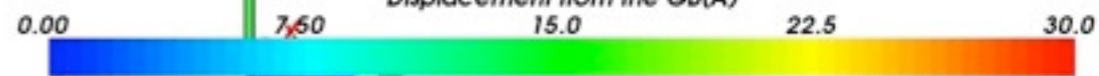
Displacement from the GB(A)



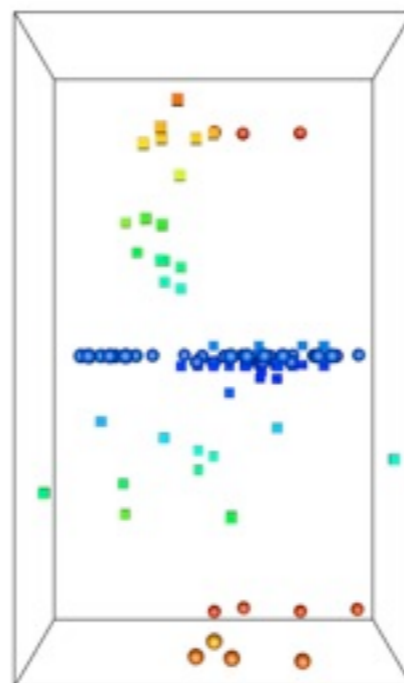
350.00 ms
8470 Atoms
88 Visible
43 Ni Vacancies
45 Ni Interstitials



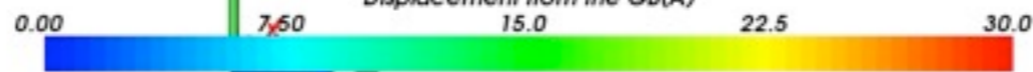
Displacement from the GB(A)



23.02 ms
8470 Atoms
95 Visible
49 Ni Vacancies
46 Ni Interstitials

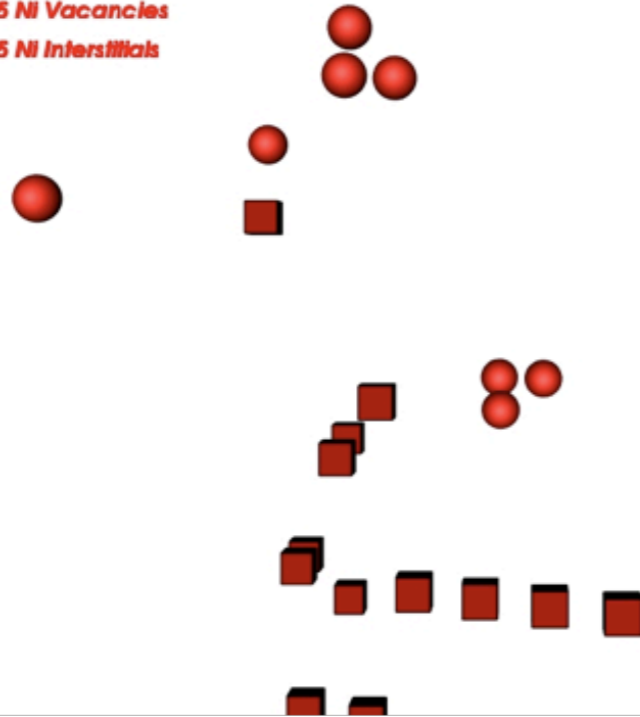


Displacement from the GB(A)

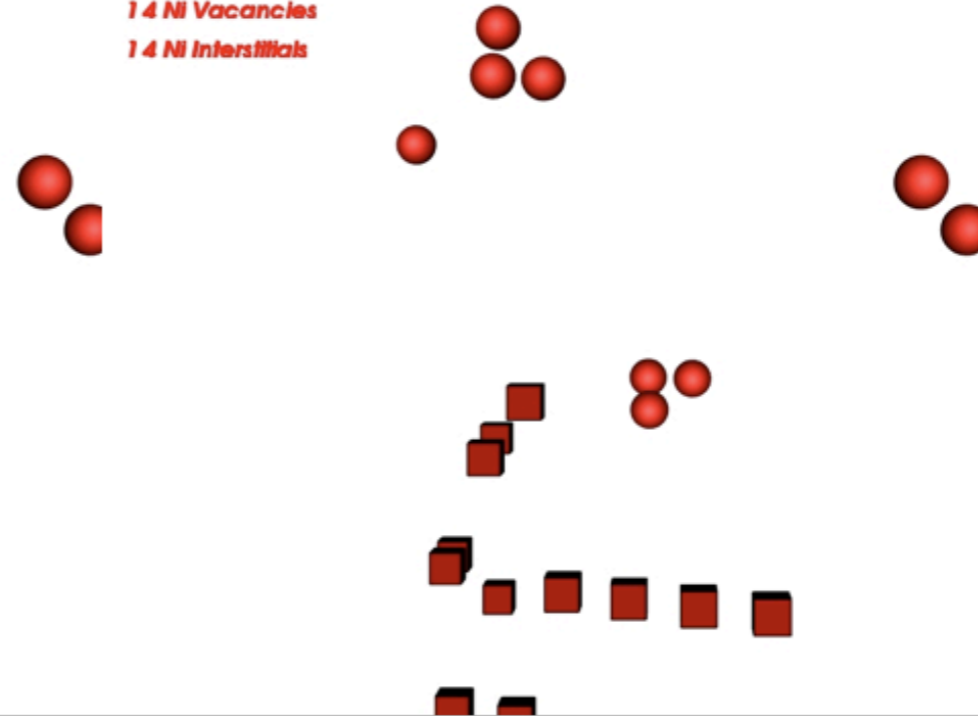


Concerted Events

220.53 ps
56700 Atoms
30 Visible
15 Ni Vacancies
15 Ni Interstitials



227.02 ps
56700 Atoms
28 Visible
14 Ni Vacancies
14 Ni Interstitials



227.02 ps
56700 Atoms
3 Visible
3 Ni



Challenges

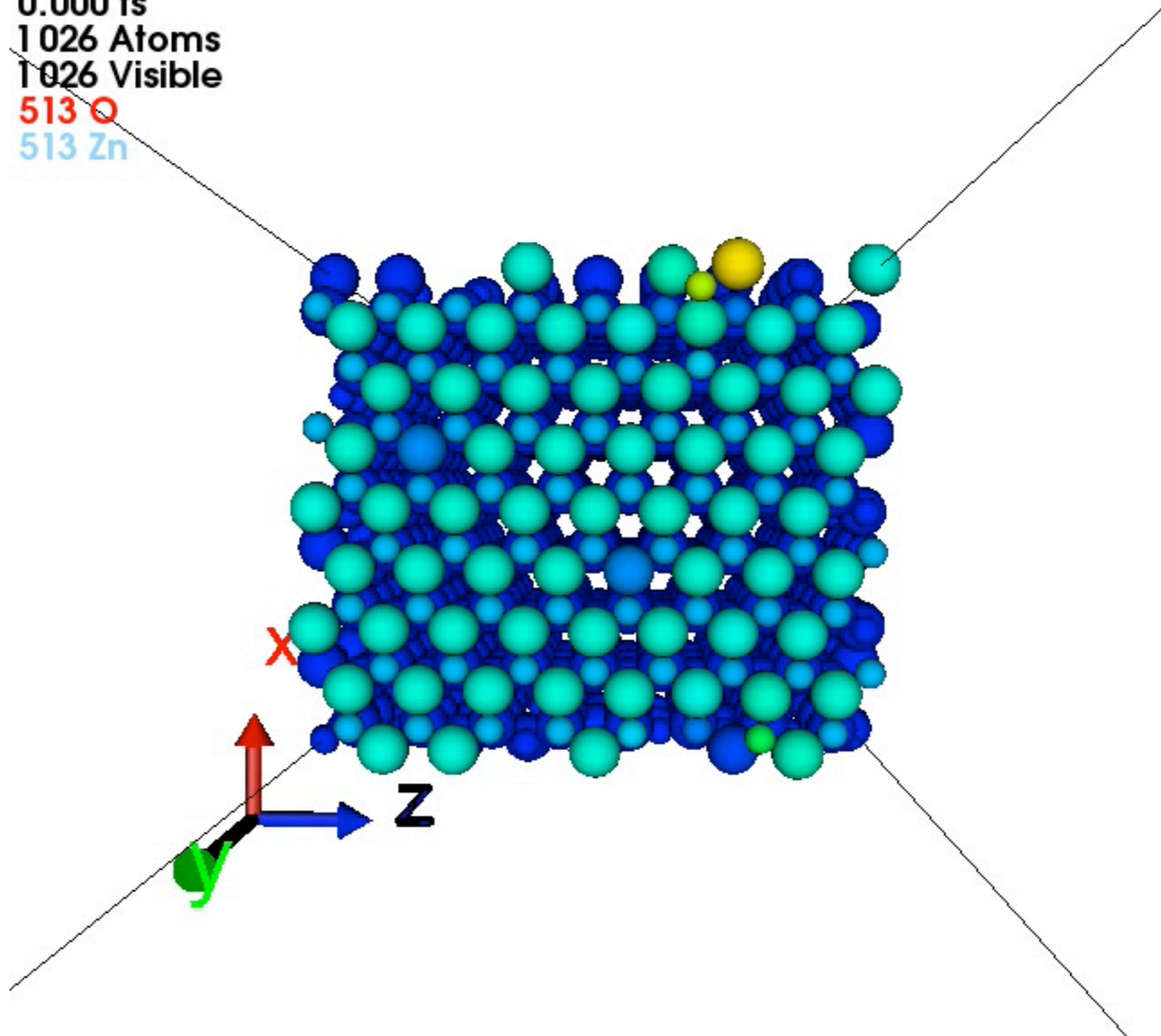
- Speed
- Fidelity

Speed

- Main cost is function evaluations - 98-99% of time.
- Need to minimise by choice of methodology and parameters.
- Filtering out low energy transitions that don't contribute to system evolution.
- Issue in most systems we have studied.

Challenges

0.000 fs
1 026 Atoms
1 026 Visible
513 O
513 Zn



Fidelity

- How many searches is enough?
- Fixed number of searches checked to see whether sufficient.

Conclusions

- Very powerful tool for studying evolution of systems.
- Can already study real systems.
- Still work to be done on refining the methodology.
- Developments needed:
 - Auto-identification of events that don't contribute to the system evolution.
 - Understanding when to stop searches.

Acknowledgements

- Roger Smith
- Chris Scott, Sabrina Blackwell, Zainab Al Tooq, Tomas Lazusaskas, Louis Vernon, Marc Robinson, Ed Sanville, Stewart Gordon, Hurry Hurchand
- AWE, EPSRC, EU